

$$\dot{Q}_{\text{loss}} = \dot{Q}_{\text{rad}} + \dot{Q}_{\text{conv}}$$

$$=$$

$$\dot{Q}_{\text{rad}} = \epsilon \sigma T_s^4 = 15.67 \cdot 10^{-8} \frac{\text{W}}{\text{m}^2 \cdot \text{K}^4} \cdot (548\text{K})^4$$

$$\dot{Q}_{\text{rad}} = 5113 \text{ W/m}^2 \cdot \frac{\pi D^2}{4} = 15.7 \text{ W}$$

$$\dot{Q}_{\text{conv}} = \text{natural convection from flat plate} = \bar{h} A (T_s - T_{\infty})$$

$$\bar{h} \text{ from Nusselt \# (table 9-1)}$$

$$L = A_s / P = \frac{\pi D^2}{4 \pi D} = \frac{D}{4}$$

$$= \frac{25 \text{ cm}}{4} = 0.0625 \text{ m}$$

$$Ra = \frac{g \beta (T_s - T_{\infty}) L_c^3}{\nu \alpha}$$

fluid properties at $T_{f,lm} = 147.5^{\circ}\text{C} = 420.5 \text{ K}$

$$k = 0.034 \text{ W/mK}$$

$$\beta = \frac{1}{T} = 0.00238 \text{ 1/K} \quad \textcircled{B}$$

$$\nu = 2.831 \cdot 10^{-6} \text{ m}^2/\text{s} \quad \textcircled{A}$$

$$\alpha = 4.026 \cdot 10^{-5} \text{ m}^2/\text{s} \quad \textcircled{A}$$

$$Ra = \frac{9.8 \text{ m/s}^2 \cdot 0.00238 \frac{1}{\text{K}} (255 \text{ K}) (0.0625 \text{ m})^3}{2.831 \cdot 10^{-6} \text{ m}^2/\text{s} \cdot 4.026 \cdot 10^{-5} \text{ m}^2/\text{s}}$$

$$= 1.27 \cdot 10^8, \text{ so}$$

$$Nu_L = 0.15 Ra_L^{1/3}$$

$$= 75$$

$$\bar{h} = \frac{Nu \cdot k}{L} = 41.1 \text{ W/m}^2 \cdot \text{K}$$

$$\dot{Q}_{\text{conv}} = 41.1 \frac{\text{W}}{\text{m}^2 \cdot \text{K}} \cdot \frac{\pi D^2}{4} \cdot (T_s - T_{\infty})$$

(0.0625 m)²
255 K

$$\dot{Q}_{\text{conv}} = 32 \text{ W}$$

$$\dot{Q}_{\text{total}} = \dot{Q}_{\text{conv}} + \dot{Q}_{\text{netrad}}$$

$$= 32 \text{ W} + 15.7 \text{ W}$$

$$= 47.7 \text{ W}$$

2.

a. To what minimum temperature must a filament be heated so 10% of energy is visible ($\lambda < 0.8 \mu\text{m}$)

assume: black or grey body so f_λ (blackbody fcs) apply.

for $\lambda T > 2200 \mu\text{mK}$, $f_\lambda > 0.10$ (10%)

$$\text{if } \lambda < 0.8 \mu\text{m} \rightarrow T \geq \frac{2200 \mu\text{mK}}{0.8 \mu\text{m}} = \boxed{2750\text{K}}$$

b. at what wavelength will emission be maximum?

Wien's law $\lambda_{\text{max}} T = 2898 \mu\text{mK}$

$$\lambda_{\text{max}} = \frac{2898 \mu\text{mK}}{2750\text{K}} = \boxed{1.053 \mu\text{m}}$$

↓

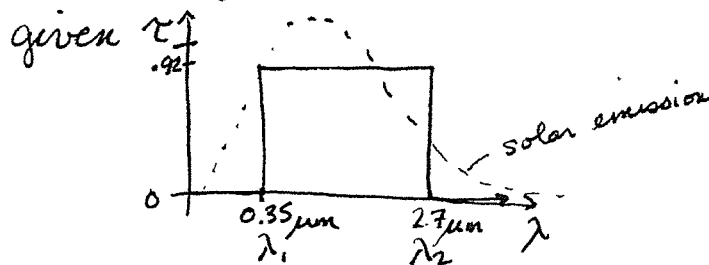
Infrared

↑ max emission as heat, not visible light!

3.

Glass transmits 92% of radiation in λ range $0.35 - 2.7 \mu\text{m}$. It is opaque otherwise.

What % of solar radiation will it transmit?



assume:
sun is bb radiator at 5800 K.

$$\lambda_1 T_{\text{sun}} = 0.35 \cdot 5800 = 2030 \mu\text{mK}$$

$$\lambda_2 T_{\text{sun}} = 2.7 \cdot 5800 = 15660 \mu\text{mK}$$

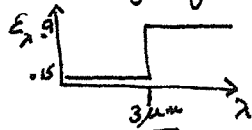
step 1: what fraction of sunlight is between 0.35 & $2.7 \mu\text{m}$?

$$f_{\lambda_1 T} - f_{\lambda_2 T} = f_{0.2 T} - f_{0.7 T} = 0.97 - 0.07 = 0.90$$

step 2: 92% of this light is transmitted: % transmitted =

$$\tau \cdot f_{0.35-2.7} = 0.92 \cdot 0.9 = 0.83 \text{ or } \boxed{83\%}$$

4. The emissivity of a surface is 0.15 for $\lambda < 3\mu\text{m}$
 & 0.9 for $\lambda > 3\mu\text{m}$



- a) what is $\bar{\epsilon}$ for $T_{\text{surface}} = 5800\text{ K}$?

$$\begin{aligned}\bar{\epsilon} &= \epsilon_{0 \rightarrow 3\mu\text{m}} \cdot f_{0 \rightarrow 3\mu\text{m}} + \epsilon_{3\mu\text{m} \rightarrow \infty} f_{3\mu\text{m} \rightarrow \infty} \\ &= 0.15 \cdot f_{0 \rightarrow 3\mu\text{m}} + 0.9 \cdot (1 - f_{0 \rightarrow 3\mu\text{m}})\end{aligned}$$

$$\lambda T \text{ for } \lambda = 3\mu\text{m}, T = 5800\text{ K} \rightarrow \lambda T = 17400\mu\text{m K}$$

$$f_{17400} = 0.979$$

$$\bar{\epsilon}_{5800\text{ K}} = 0.15(0.979) + 0.9(0.021)$$

$$\boxed{\bar{\epsilon}_{5800\text{ K}} = 0.166}$$

- b) same for $\bar{\epsilon}_{300\text{ K}}$ $\Rightarrow$

$$\lambda T = (3\mu\text{m}) \cdot (300\text{ K}) = 900\mu\text{m K}$$

$$f_{900\mu\text{m K}} = 0.000169$$

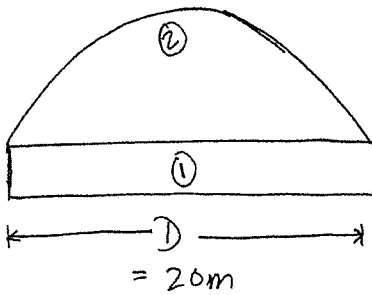
$$\bar{\epsilon}_{300\text{ K}} = 0.15 \cdot (0.000169) + .9(.999832)$$

$$\boxed{\bar{\epsilon}_{300\text{ K}} = 0.900}$$

- c) the absorptivity of this surface to sunlight is equal to its emissivity for a ^{spectral} profile at 5800K
 so $\boxed{\alpha_{\text{solar}} = 0.166}$

- d) this surface is a better solar reflector. It only absorbs 16% of energy at solar wavelengths, but it emits 90% of possible energy at room temp.

5.



a)

$$F_{1 \rightarrow 1} + F_{1 \rightarrow 2} = 1$$

$$F_{2 \rightarrow 1} + F_{2 \rightarrow 2} = 1$$

$$F_{2 \rightarrow 2} = 1 - F_{2 \rightarrow 1}$$

$$A_2 F_{2 \rightarrow 2} = A_2 (1 - F_{2 \rightarrow 1})$$

$$A_2 F_{2 \rightarrow 2} = A_2 - A_2 F_{2 \rightarrow 1}$$

$$A_2 F_{2 \rightarrow 2} = A_2 - A_1 F_{1 \rightarrow 2}$$

$$F_{2 \rightarrow 2} = \frac{A_2 - A_1 F_{1 \rightarrow 2}}{A_2}$$

$$A_1 = \pi D^2 / 4 = 314 \text{ m}^2$$

$$A_2 = \frac{1}{2} 4\pi (D/2)^2 = 628 \text{ m}^2$$

$$F_{2 \rightarrow 2} = 1 - \frac{A_1}{A_2} = 1 - \frac{314}{628}$$

$$F_{2 \rightarrow 2} = \frac{1}{2}$$

b)

$$\dot{Q}_{\text{net}, 2 \rightarrow 1} = A_2 \sigma F_{2 \rightarrow 1} (T_2^4 - T_1^4)$$

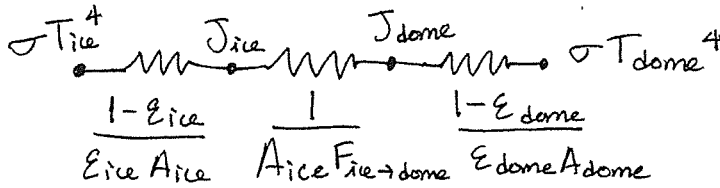
$$= A_1 F_{1 \rightarrow 2} \sigma (T_2^4 - T_1^4)$$

$$= (314 \text{ m}^2) (5.67 \times 10^{-8} \text{ W/m}^2 \text{ K}^4) (288 \text{ K}^4 - 268 \text{ K}^4)$$

$$\dot{Q}_{\text{net}, 2 \rightarrow 1} = 30.7 \text{ kW}$$

c) $\epsilon_{ice} = 0.9$

$\epsilon_{dome} = 0.1$



$$\dot{Q}_{dome \rightarrow ice} = \frac{\sigma (T_{dome}^4 - T_{ice}^4)}{R_{total}}$$

$$R_{total} = \frac{1 - \epsilon_{ice}}{\epsilon_{ice} A_{ice}} + \frac{1}{A_{ice} F_{ice \rightarrow dome}} + \frac{1 - \epsilon_{dome}}{\epsilon_{dome} A_{dome}}$$

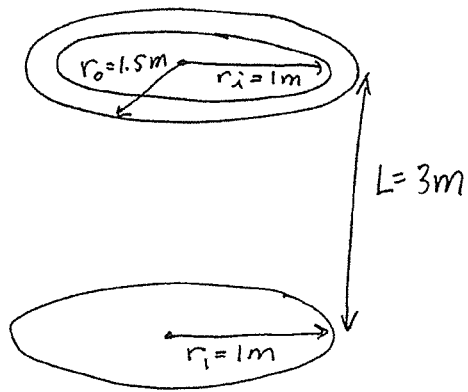
$$= \frac{1 - 0.9}{0.9(314)} + \frac{1}{(314)(1)} + \frac{1 - 0.1}{0.1(628)}$$

$$= 0.0179$$

$$\dot{Q}_{dome \rightarrow ice} = \frac{(5.67 \times 10^{-8})(288^4 - 268^4)}{0.0179}$$

$$\dot{Q}_{dome \rightarrow ice} = 5.46 \text{ kW}$$

6.



$$F_{\text{disk} \rightarrow \text{ring}} = F_{\text{disk} \rightarrow r_o} - F_{\text{disk} \rightarrow r_i}$$

$$F_{\text{disk} \rightarrow r_o} = \frac{1}{2} \left\{ S - \left[S^2 - 4 \left(\frac{r_o}{r_1} \right)^2 \right]^{1/2} \right\}$$

$$S = 1 + \frac{1 + R_o^2}{R_i^2}, \quad R_o = r_o/L, \quad R_i = r_i/L$$

$$F_{\text{disk} \rightarrow r_i} = 1 + \frac{1 - \sqrt{4R^2 + 1}}{2R^2}, \quad R = \frac{r_i}{L} = \frac{r_1}{L}$$

$$F_{\text{disk} \rightarrow \text{ring}} = \frac{1}{2} \left\{ 1 + \frac{1 + \left(\frac{r_o}{L} \right)^2}{\left(\frac{r_1}{L} \right)^2} - \left[\left(1 + \frac{1 + \left(\frac{r_o}{L} \right)^2}{\left(\frac{r_1}{L} \right)^2} \right)^2 - 4 \left(\frac{r_o}{r_1} \right)^2 \right]^{1/2} \right\}$$

$$- \left\{ 1 + \frac{1 - \sqrt{4R^2 + 1}}{2R^2} \right\}$$